Stratospheric Influence on the Composition of the Mid- and Upper-Troposphere over North America sampled by the NASA DC-8 during INTEX A

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ABSTRACT
A comprehensive atmospheric chemistry payload was deployed on NASA’s DC-8 airborne laboratory for the INTEX A campaign in July and August, 2004. In-situ and remote measurements from the DC-8 were made to further understanding of the magnitude and distribution of anthropogenic emissions in North America, the processes exporting these emissions from the continental boundary layer, and the transformations they underwent as they were transported downwind to eastern North America and the Atlantic Ocean. We present the vertical and geographic distributions, and relationships between, $^7$Be, O$_3$, and HNO$_3$ in the INTEX A data set. Our objective is to identify stratospherically impacted air masses in the upper troposphere, and to estimate their influence on the composition of the upper troposphere, since unrecognized stratospheric influence may confound efforts to understand the fate of surface emissions lofted to the upper troposphere. Beryllium-7 and the HNO$_3$/O$_3$ ratio identify similar distinct stratospheric and tropospheric populations in the upper troposphere. Stratospherically impacted air was most common over the north eastern portion of the INTEX A study area. Estimates of the impact of stratospheric contributions on the upper tropospheric
budgets of O$_3$ and HNO$_3$ based on $^7$Be suggest that the stratosphere was the major source of both gases above 6 km. Using the HNO$_3$/O$_3$ ratio to filter out stratospheric influence suggest that the stratospheric source was important for HNO$_3$ (but much weaker than implied by $^7$Be) and minor for O$_3$. We investigate possible causes for these divergent results, but conclude that available data does not allow us to constrain the impact of the stratosphere on upper tropospheric composition in the INTEX A domain with useful precision using these tracers.

INTRODUCTION

The Intercontinental Chemical Transport Experiment, Phase A (INTEX A) was the NASA Tropospheric Chemistry Program’s contribution to the International Consortium for Atmospheric Research on Transport and Transformation (ICARTT) intensive campaign in summer 2004. The overall campaign involved many different research assets deployed to sample the troposphere above North America and the North Atlantic to improve understanding of the magnitude, transformations, and fate of anthropogenic emissions from North America. INTEX had multiple science objectives of particular interest to NASA, in addition to those specific to ICARTT, these are detailed by Singh et al. [this issue a]. This overview of INTEX also provides specific objectives of each sortie that was flown, and describes the payload on board the DC-8 airborne laboratory for this mission.

One of the shared objectives of INTEX A and ICARTT 2004 was improved understanding of the export of anthropogenic and natural emissions from the North American boundary layer into the free troposphere, and subsequent transport over the North Atlantic and onto Europe. Very long-range transport of surface-based emissions is often facilitated if they are lofted into the upper troposphere where winds are usually stronger. Several companion papers discuss the evidence for significant vertical transport of North American emissions out of the boundary layer and lower troposphere by numerous convective storms over the continent during the campaign [e.g., Bertram et al., 2006; Cooper et al., this issue]. Polluted boundary layer air can be modified by loss of soluble components in wet convection, the addition of NO$_x$ from lightning in the convective systems, and rapid photochemical processing within the upper troposphere.
An additional factor complicating efforts to understand the transformations and fate of anthropogenic emissions that reach the upper troposphere is mixing with stratospherically influenced air, which has enhanced mixing ratios of O₃ and HNO₃, two trace gases that are pollutants of concern in the lower troposphere. Browell et al. [this issue] identify and tabulate airmasses with stratospheric influence, determined by UV-DIAL, above and below the DC-8 on all INTEX A flights and Avery et al. [this issue] characterize the dynamical processes responsible for bringing stratospheric air into the troposphere in the INTEX A study region. While it is generally possible to recognize upper tropospheric air with a large stratospheric influence (and filter these data point out of some analyses, or track them as a special class of upper tropospheric air masses, e.g., Singh et al., [this issue b]), there are often many more upper tropospheric air masses with weak stratospheric characteristics that are more difficult to identify.

In this paper we use the distributions of ⁷Be (a cosmogenic radionuclide formed mainly in the stratosphere), HNO₃, and O₃, and their inter-relationships, to: 1) identify air masses with both strong and weak stratospheric influences in the upper troposphere (above 6 km), and 2) attempt to constrain the impact of stratosphere to troposphere exchange on the upper tropospheric budgets of O₃ and HNO₃ above North America during INTEX A.

METHODS

The NASA DC-8 flew 18 science missions during INTEX A, starting and ending at NASA, Dryden. Two field bases were utilized, Middle America, just east of St. Louis, MO, and Pease Trade Port, in Portsmouth, NH. Local flights from Middle America were mainly over continental North America, but did extend to the Gulf of Mexico. Sorties from Pease covered much of the eastern US, but emphasized areas over extreme eastern Canada and the North Atlantic where chemical transport models indicated much of the export of North American emissions was occurring (see Singh et al., [this issue a]).

In situ measurements of HNO₃ and O₃ were made on board the DC-8 by mist chamber/ion chromatography (MC/IC) and chemiluminescence (CL), respectively. The HNO₃ and O₃ systems have extensive track records in previous airborne campaigns on the DC-8 and other platforms. The automated MC/IC system was described by Scheuer et al. [2003], continued improvements to the IC analytical system allowed us to use a
sample integration time of 106 seconds throughout INTEX A. The FASTOZ CL instrument flown on recent DC-8 campaigns is described in Avery et al. [2006], O$_3$ measurements are reported at one second intervals. This paper relies on a merged data set produced at NASA Langley, wherein the faster O$_3$ measurements are averaged over the MC sample collection interval.

The activity of $^7$Be in bulk aerosol samples captured on glass fiber filters was determined by gamma spectroscopy, as described in Dibb et al. [1996]. We used the dual inlet aerosol probe described by McNaughton et al. [2006] and Dibb et al. [2003a] to expose simultaneous filter samples for the radionuclide tracers and for soluble ions. The aerosol-associated ion data is discussed by Clarke et al. [this issue], Tang et al. [this issue] and Thornhill et al. [this issue], this paper uses $^7$Be as an indicator of stratospheric influence.

Aerosol samples are collected only on level flight legs, so that sample flow rate can be adjusted to maintain isokinetic flow through the inlet nozzle. Filters were exposed for less than 5 minute intervals below 3 km pressure altitude, but integration time increased at higher altitudes. The overall average filter exposure interval was 7.1 minutes (range 2 – 27 minutes) during INTEX A, increasing to 11.1 minutes (range 5 – 27) above 6 km. Because one of the INTEX A objectives was to profile the entire troposphere above and downwind of North America [Singh et al., this issue a], a large portion of each flight was spent climbing and descending (e.g., not sampling aerosols) (Figure 1). As a result, 45% of the MC sampling intervals do not have corresponding data on filter-collected aerosols. In the upper troposphere (> 6 km) the fraction of MC sampling time without $^7$Be data decreases to 27%, however, the average filter sample overlaps with approximately 6 MC samples. Rather than averaging the HNO$_3$, O$_3$, and NO$_2$ mixing ratios over the filter integration times, we repeat the $^7$Be activity within the corresponding MC sample intervals for subsequent analysis of the relationships between $^7$Be and the trace gases.
RESULTS

Beryllium-7 as Stratospheric Tracer.

Profiles of $^7\text{Be}$ and O$_3$ show gradual increases with altitude, and numerous large enhancements in individual samples above 6 km (Figure 2). Avery et al. [this issue] examine several of the DC-8 encounters with stratospheric air below the tropopause during INTEX A, and find that the cross-tropopause exchange had occurred both within the INTEX A region, as well as upwind, in different cases. They show that the subtropical jet largely determined the location of stratosphere-to-troposphere exchange (STE) impacting the troposphere above North America during INTEX A. The steady increase with altitude is consistent with STE constituting an important source of both $^7\text{Be}$ and O$_3$ in the upper troposphere (UT) over North America during INTEX A. Very large enhancements in both species at the highest altitudes reflect penetration of the lower stratosphere or tropopause folds, but finding $^7\text{Be} \gg 1000$ fci m$^{-3}$ STP and O$_3 \gg 100$ ppb as low as 6 km is surprising. The samples with very low $^7\text{Be}$ and O$_3$ in the UT are also of note, as they reflect convective pumping of boundary layer air into the UT [see Bertram et al., 2006].

Nitric acid mixing ratios peak in the polluted continental boundary layer, so, in the context of a profile through the altitude sampling range of the DC-8 during INTEX A, the UT looks relatively clean (Figure 2). However, the mean mixing ratios in the 5-8 km, and > 8 km, altitude ranges (264 ppt and 232 ppt, respectively) are elevated in comparison to what we have observed in the UT on previous campaigns. For example, the mean HNO$_3$ mixing ratio above 6 km over the Pacific during PEM West B was 130 ppt (when 6 samples in a tropopause fold near Japan were excluded) [Dibb et al., 1997]. During TRACE P the mean mixing ratio was 175 ppt of HNO$_3$, in the same altitude bin over a similar North Pacific region [Dibb et al., 2003b]. Most of the INTEX A MC samples above 6 km with HNO$_3$ in excess of 500 ppt also had elevated (> 150 ppb) O$_3$ mixing ratios. Enhanced UT HNO$_3$, correlated with elevated O$_3$, would also be consistent with STE impacting the composition of the UT over North America during the INTEX A campaign. However, it is also possible for tropospheric photochemistry to produce both O$_3$ and HNO$_3$ [e.g., Bertram et al., 2006], accounting for some of their
observed upper UT burdens, and making positive correlations between them an ambiguous indicator of stratospheric origin.

Relatively weak tropospheric production of $^7$Be (globally 2/3 of production is in the stratosphere [Lal et al., 1958]) and rapid removal of the aerosols it is associated with by precipitation scavenging prevent purely tropospheric activities from exceeding several 100 fCi m$^{-3}$. Kritz et al. [1991] suggest that 500 fCi m$^{-3}$ STP is a good estimate for UT $^7$Be activity, but this includes a stratospheric contribution. In contrast, in the stratosphere the sinks of $^7$Be are radioactive decay and STE, leading to lifetimes on the order of a year compared to about a month in the upper troposphere [e.g., Dibb et al., 1994]. At 15 km altitude (close to the zone of maximum production) near 40$^\circ$ the steady state $^7$Be activity has been estimated to be 8 pCi m$^{-3}$ [Bhandari et al., 1966], thus STE causes large, and unambiguous, enhancements in $^7$Be below the tropopause. Correlations of other species with $^7$Be in the troposphere can therefore be used to assess whether STE is also a significant source for them.

In the INTEX A data set O$_3$ is reasonably correlated with $^7$Be when all samples are considered (Figure 3, solid line in upper panel). For HNO$_3$ there are clearly two separate populations, a tropospheric group with high HNO$_3$ but low $^7$Be, and a stratospheric branch where they increase together. We selected 600 fCi $^7$Be m$^{-3}$ STP as an admittedly arbitrary, but reasonable, threshold to define lower stratospheric and strongly STE impacted UT air masses in the INTEX A data set, and present the linear relation between O$_3$ and $^7$Be in these “stratospheric” samples in the lower panel of Figure 3. Since the HNO$_3$ data shows there is clearly a distinct stratospheric and stratospherically impacted subset of the INTEX A data set, we also used the 600 fCi $^7$Be m$^{-3}$ STP threshold to examine the O$_3$ – $^7$Be relationship in this group of samples with strong stratospheric influence. The linear regression of O$_3$ against $^7$Be > 600 fCi m$^{-3}$ STP (shown as the dotted line in upper panel of Figure 3) has a 27% lower intercept and a 13% steeper slope than the fit to the entire data set. The fact that the two relationships are so similar reflects the dominant impact of the small number of samples with very strong stratospheric character, over the cloud of points with low $^7$Be activity and low to moderate O$_3$ mixing ratios, on the regression.
If we assume that the regression lines in Figure 3 represent mixing lines between tropospheric air and a stratospheric source with constant, greatly enhanced, $^7$Be, O$_3$, and HNO$_3$, we can use measured $^7$Be to estimate the stratospheric mixing ratios of O$_3$ and HNO$_3$ in all INTEX A samples for which we have $^7$Be data [e.g., Dibb et al., 2000, 2003c]. An estimate of the tropospheric fraction of O$_3$ and HNO$_3$ in each sample can then be obtained by subtracting the calculated “stratospheric” mixing ratio from the actual measurement (i.e., Trop O$_3$ = Measured O$_3$ – estimated Strat O$_3$). Using the relationship for the full data set (Strat O$_3$ = (0.039 x $^7$Be) + 50.1, from the solid line in upper panel of Figure 3) yields a mean tropospheric O$_3$ mixing ratio of 8.8 ppb above 8 km, and –2.2 ppb between 5 and 8 km. If the regression using the $^7$Be > 600 fCi m$^{-3}$ STP threshold is used (dotted line in upper panel of Figure 3), the estimated mean tropospheric O$_3$ is 8.6 and 17.7 ppb, respectively, in the same altitude bins.

Given observed mean O$_3$ mixing ratios of 91 and 71 ppb in these altitude bins, this analysis suggests that more than 90% of the O$_3$ above 8 km came from the stratosphere, and 75% to more than 100% of O$_3$ in the mid troposphere is also stratospherically derived. Parallel calculations for HNO$_3$ suggest an 88% contribution from the stratosphere above 8 km (28 ppt mean tropospheric HNO$_3$ compared to 232 ppt observed mean) and a 61% contribution in the 5 – 8 km range (104 ppt versus 264 ppt). In contrast, using a filter to screen out UT samples with obvious stratospheric influence (all samples above 7 km with O$_3$ > 120 ppb and H$_2$O vapor < 100 ppm) results in mean mixing ratios of O$_3$ and HNO$_3$ above 8 km of 79 ppb and 203 ppt, respectively [Singh et al., this issue b]. This approach suggests that STE makes about a 13% contribution to the mean mixing ratios of both gases in the INTEX upper UT. However, it excludes, by definition, the possibility of any stratospheric component in an air mass with less than 120 ppb O$_3$, and would consider the samples with clear stratospheric signatures of O$_3$ and $^7$Be between 6 and 7 km (Figure 2) to be tropospheric. The large number of samples with O$_3$ mixing ratios in the ~ 60 to 120 ppb range with 500 to 1000 ppt of HNO$_3$ (Figure 4) create ambiguity when using threshold values of any tracer, or combination of tracers, to screen out stratospheric influence in the troposphere.

*Pierce et al.* [this issue] use the RAQMS 3-D chemical transport model, with chemical data assimilation, to assess the tropospheric budgets of O$_3$ and HNO$_3$ over
North America during INTEX A. They suggest that the net cross-tropopause flux of both gases within the CONUS domain is actually upward into the lower stratosphere. However, they also find that much of the UT O$_3$ and HNO$_3$ advected into the INTEX A domain across the western border within the UT, after having been injected into the UT by STE upwind over the Pacific. This may be consistent with the $^7$Be-based estimates of very strong stratospheric influence on the UT during INTEX A, since the measured $^7$Be activity provides no insight into where STE may have occurred. Analysis of ozone sondes launched in support of the ICARTT campaign also lend support to the assertion of significant stratospheric influence in the North American troposphere during INTEX A [Thompson et al., this issue], as does the analysis of O$_3$ curtains derived from the UV-DIAL system on the DC-8 [Browell et al., this issue].

The estimates of stratospheric influence on the composition of the upper troposphere over North America from $^7$Be would appear to preclude significant photochemical production of O$_3$ and HNO$_3$ above 6 km during INTEX A. This assertion is at odds with several model studies of the INTEX A data set. Bertram et al. [2006] simulate rapid production of both O$_3$ (up to 7.5 ppb d$^{-1}$ for the first two days) and HNO$_3$ in air masses recently convected into the UT. They also estimate that 40% of the air masses sampled by the DC-8 between 7.5 and 11.5 km had been convectively uplifted within less than two days before sampling. It should be noted that the recently convected airmasses in the UT identified by elevated NO$_x$/HNO$_3$ (or NO$_2$/HNO$_3$) start with low mixing ratios of O$_3$ (reflecting origin in the boundary layer) and HNO$_3$ (which is effectively scavenged in the convective systems). Cooper et al. [this issue] also suggest that convected boundary layer NO$_x$, and NO$_x$ from lightning in the convective systems, make a major contribution to O$_3$ enhancements observed in the UT during INTEX A, particularly over the southern regions. Constrained box model calculations find net production of O$_3$ across the INTEX A UT as well [Crawford et al., this issue]. The RAQMS simulations discussed above also find significant net O$_3$ production in the troposphere, but mainly in the first few weeks of INTEX A, with net loss apparent in the tropospheric column after the end of July [Pierce et al., this issue].

It should be borne in mind that requiring observations of $^7$Be above detection limit for our analysis reduces the data sets from 1464 samples to 828 samples above 8
km, and from 1075 to 395 samples between 5 and 8 km. We also acknowledge that the estimates of stratospheric mixing ratios have large uncertainty due to the significant scatter about the regression (Figure 3). Estimating the tropospheric residual has additional uncertainty due simply to the precision of the multiple measurements needed to make the calculation. We will return to these, and other, issues impacting this approach to estimating stratospheric influence on the troposphere in the discussion.

**HNO\textsubscript{3}/O\textsubscript{3} as a Stratospheric Tracer.**

Given that O\textsubscript{3} and HNO\textsubscript{3} are each correlated with \(^7\)Be in the upper troposphere, a positive correlation between O\textsubscript{3} and HNO\textsubscript{3} would be expected. A scatter plot reveals that there are two distinct populations in the complete INTEX A data set (Figure 4, upper panel). Samples with > 1 ppb HNO\textsubscript{3} but O\textsubscript{3} generally < 100 ppb define the polluted boundary layer and lower troposphere, with a smaller stratospherically impacted group of samples defined by a much shallower HNO\textsubscript{3}/O\textsubscript{3} slope, but higher O\textsubscript{3} mixing ratios. The INTEX A data points are plotted on top of the results we obtained from the DC-8 during the Polar Aura Validation Experiment (PAVE) in January, 2005 [Dibb et al., 2006]. The DC-8 was also based at Pease for PAVE and conducted 6 sorties north into the Arctic; the vast majority of sampling was conducted in the lower stratosphere. A very small tropospheric subset of PAVE results is present, but not visible, under the INTEX A data. The INTEX A stratospherically influenced samples fall along the bottom of the distribution of PAVE stratospheric samples. During PAVE, the HNO\textsubscript{3}/O\textsubscript{3} ratio was always near 3.5 ppt/ppb in the stratosphere, except when perturbed by renitrification at DC-8 flight levels [Dibb et al., 2006]. The PAVE stratospheric ratio is also shown in the lower panel of Figure 4, superimposed on all INTEX A samples collected above 6 km. The 3.5 ppt/ppb slope is a reasonable depiction of the INTEX A stratospherically impacted population, if the line is extrapolated to 40 ppb O\textsubscript{3} at 0 ppt HNO\textsubscript{3}.

We observed that the HNO\textsubscript{3}/O\textsubscript{3} ratio above 6 km during INTEX A was often near the PAVE stratospheric value of 3.5 ppt/ppb, except when HNO\textsubscript{3} was depressed in recently convected air masses, as indicated by simultaneous enhancements of NO\textsubscript{2}/HNO\textsubscript{3} and decreases in HNO\textsubscript{3}/O\textsubscript{3} (Figure 5), and more rarely when HNO\textsubscript{3} was enhanced in biomass burning plumes (not observed during the example Flight #11 shown in Figure 5).
Scatter plots of HNO$_3$ and O$_3$ against the HNO$_3$/O$_3$ ratio also show two subpopulations in the INTEX A UT data set, with the stratospherically impacted group identified in Figures 3 and 4 largely restricted to values of HNO$_3$/O$_3$ between 2.0 and 4.0 (Figure 6). This suggest that the HNO$_3$/O$_3$ ratio maybe complementary to $^7$Be as a tracer of stratospheric influence in the UT. If true, this ratio is available at the 106 second MC sample resolution throughout the UT.

All INTEX A UT (>6 km) samples were divided into 4 geographic bins: west of 80 W and south of 41 N (SW), west of 80 W and north of 41 N (NW), east of 80 W and south of 41 N (SE), and east of 80 W and north of 41 N (NE). The spatial bins were also divided into 3 upper tropospheric altitude bins and the fraction of all MC sampling intervals with HNO$_3$/O$_3$ near the PAVE stratospheric ratio (2.0–4.0 ppt/ppb (Figure 6)) were compiled (Table 1). The fraction of UT samples identified as stratospherically impacted by this ratio was highest in the NE bin, followed by the SW. One might have expected the NW bin to also show frequent stratospheric influence, reflecting the lower tropopause at higher latitudes. Browell et al. [this issue] also find more stratospheric influence in the troposphere in the north east portion of the INTEX A study area. Enhanced STE over the NE bin likely reflects regionally lower tropopause heights associated with the persistent trough over eastern North America during INTEX A that largely overlapped the period when the DC-8 was based in New Hampshire [Fuelberg et al., this issue].

In the NE and SW regions, the frequency of samples with 2.0 < HNO$_3$/O$_3$ < 4.0 is highest in the top altitude range and decreases downward. This is consistent with STE being the source of the air masses with this proposed characteristic ratio. The decreasing trend in apparent stratospheric influence as altitude increases seen in the NW and SE regions is less easy to understand, especially the high fractions in the 6–8 km range (Table 1). It may be that many of the low altitude samples with the “stratospheric” HNO$_3$/O$_3$ signature in these regions come from the intersection of the two arms of the distributions in Figure 6.

If the HNO$_3$/O$_3$ ratio is reflecting stratospheric influence, weighting the percentages in Table 1 by the number of samples collected in each of the bins suggests that 34% of air masses above 8 km across the entire INTEX A study region were
impacted by STE. Fuelberg et al. [this issue] find that 27% of back trajectories from the DC-8 flight tracks at all altitudes during INTEX A encountered the stratosphere within the previous 10 days. Presumably that percentage is higher at higher altitudes, so our estimate of 34% of samples above 8 km having stratospheric influence seems reasonable, and may even be conservative. If this is combined with the Bertram et al. [2006] estimate that 40% of air sampled between 7.5 and 11.5 km was lifted from the boundary layer within two days, the UT would seem to be dominated by air mixing in from other regions of the atmosphere.

We are suggesting that the HNO$_3$/O$_3$ ratio identifies samples that retain a stratospheric signature, but do not necessarily have mixing ratios characteristic of the stratosphere. Similarly, recently convected air masses start with compositions reflecting their boundary layer origin (except that soluble gases and particles can be depleted by precipitation scavenging enroute to the upper troposphere), but these are quickly modified by photochemical processes Bertram et al. [2006]. Hence, it is clearly an oversimplification to expect that the composition of the upper troposphere should be a simple mixture of the three components; boundary layer air, lower stratosphere, and background upper troposphere (if such a component can be defined).

In a first order attempt to quantify the impact of stratospheric and convective inputs on UT composition we have filtered two groups of samples out of the INTEX A UT data set. The stratospherically impacted subset is defined as above, 2.0 < HNO$_3$/O$_3$ < 4.0. To select recently convected airmasses we used NO$_2$/HNO$_3$ > 1.0, which is a more restrictive criterion than that used by Bertram et al. [2006] (NO$_x$/HNO$_3$ > 1.0). These two filters identify two nearly distinct populations, just 37 of the 2251 samples above 6 km fall into both groups, most of these are in the SW region.

Mean HNO$_3$ mixing ratios above 6 km increase in nearly all of the spatial/altitude bins if recently convected air masses are excluded (Figure 7). The largest differences are seen in the > 10 km bins in all regions, with more than 17% increases in the means (up to 31% in NE). The impact of convection is also relatively strong from 8 – 10 km in the SE and NE regions, where mean HNO$_3$ increases 11 and 12%, respectively, without the convective samples. Excluding the stratospheric samples depresses mean HNO$_3$ mixing ratios in 6 of the 8 bins above 8 km (Figure 7). The decrease only exceeds 10% in the
SW > 10 km, SE 8 – 10 km, and two highest NE bins. It should be noted that the means above 8 km in NE decrease by more than 30%, with nearly equal impacts seen in the 8 – 10 and > 10 km ranges. The overall decrease above 8 km is modest compared to the 88% contribution from the stratosphere estimated from HNO$_3$ versus $^7$Be. At 6 – 8 km in all regions removing the samples identified as stratospherically impacted actually causes the HNO$_3$ mean to increase modestly, contrasting even more starkly with the estimate from $^7$Be that STE contributed 61% of the HNO$_3$ in the 5-8 km range over the INTEX A domain.

Results for O$_3$ are even more discordant with the $^7$Be based estimates. Excluding “stratospheric” samples from the UT data set decreases the mean by 15% from 8 – 10 km in the SE and NE bins, and by 20% above 10 km in NE (Figure 8). In 4 of the 5 other bins above 8 km the mean O$_3$ actually increases slightly when the stratospherically influenced samples are excluded (and the decrease in the SW region above 10 km is just over 2%). At 6 – 8 km all regions show small decreases, but the largest is < 6% (NE region). Recall that the $^7$Be-based approach attributed more than 90% of the O$_3$ above 8 km to STE. The impact of recently convected air masses on the O$_3$ content of the upper troposphere is also suggested to be very modest by this analysis (Figure 8). In 3 of the highest altitude bins, mean O$_3$ increases by ~ 5%, reaching a maximum of 15% in the NE, when the convective samples are ignored. In all other bins the means increase by < 4% (< 2% in 7 of the 8) when filtered to exclude very recent convection (Figure 8).

Excluding the samples identified as stratospheric by the HNO$_3$/O$_3$ ratio yields mean mixing ratios of 194 ppt HNO$_3$ and 82 ppb O$_3$ across the entire INTEX A study area above 8 km, corresponding to 16 and 10% reductions compared to the overall mean mixing ratios of HNO$_3$ and O$_3$ in this altitude range. Stratospheric influence above 8 km is thus estimated to be 3% larger for HNO$_3$ and 3% smaller for O$_3$ than was the case for the O$_3$ and H$_2$O vapor filter used by Singh et al. [this issue b] to remove stratospherically impacted samples from the UT data set. As noted earlier, when comparing this filter to the $^7$Be based estimates of stratospheric influence, the 120 ppb O$_3$ and 100 ppm H$_2$O criteria appear to classify significant numbers of sample with stratospheric influence as purely tropospheric and would thus be a lower bound to stratospheric influence on the composition of the UT.
DISCUSSION

There is no doubt that the DC-8 flew into the lower stratosphere on some flights, and also penetrated tropopause folds, during INTEX A (see Figures 2-4). It is also very certain that the UT contained airmasses that had been influenced by STE relatively recently, resulting in enhancements of a suite of trace gases and tracers that are greatly enriched in the stratosphere compared to the troposphere (e.g., O₃, HNO₃, and ⁷Be). Air in the lower stratosphere also has markedly lower than tropospheric concentrations of a suite of gases (e.g., H₂O, CO, CH₄, N₂O), so depletions in these gases in the UT can serve as additional tracers of STE. We examined the relationships between water vapor and the stratospheric tracers ⁷Be and HNO₃/O₃ in the INTEX A data set, but found nothing that would explain why ⁷Be implies a dominant stratospheric impact throughout the UT during INTEX A, while HNO₃/O₃ suggested a minor role for STE, particularly for O₃. Unfortunately, the availability of CO and CH₄ data is limited and no measurements of N₂O were made during INTEX A due to instrument problems.

Beryllium-7 is not photochemically produced (unlike O₃ and HNO₃), hence it has long been used as a tracer to identify stratospheric influence in the troposphere, and has less frequently been used to estimate the magnitude of such influence on the composition of the troposphere. In airborne campaigns ⁷Be has been used for both purposes [e.g., Kritz et al., 1991; Dibb et al., 2000, 2003c], but the low time resolution, and lack of data on ascending and descending flight legs, imposed by filter based sampling and analytical constraints, have been limitations. In this paper we attempt to use ⁷Be to quantify the impact of STE on the concentrations of two important trace gases in the UT, and explore the suggestion that the lower stratosphere may have a narrow characteristic range of the ratio HNO₃/O₃, such that the ratio could be a complementary tracer of stratospheric influence in the troposphere, available at much higher temporal resolution.

Enhanced ⁷Be activity and values of the HNO₃/O₃ ratio between 2.0 and 4.0 both identify a distinct subset of observations in the UT that have correlated enhancements in O₃ and HNO₃ consistent with stratospheric origin (Figures 3 and 6). In fact, the ratio (Figure 6) more clearly discriminates the stratospherically influenced O₃ observations that are apparent in the HNO₃ versus O₃ scatter plot (Figure 4) than does ⁷Be activity.
(Figure 3). Both approaches suggest that a significant number of samples in the UT during INTEX A had stratospheric influence. However, when they are used to estimate how much of the UT burden of HNO$_3$ and O$_3$ during INTEX A resulted from STE, the results are widely divergent for HNO$_3$, and even more so for O$_3$ (Figures 7 and 8). Using correlations with $^7$Be to estimate the stratospheric fraction of O$_3$ and HNO$_3$ in the UT suggests that STE is an overwhelmingly dominant influence, while the HNO$_3$/O$_3$ filter suggests a minor contribution from the stratosphere. It is unclear which is closer to the truth.

It may be that the HNO$_3$/O$_3$ ratio loses discriminatory power as the fraction of stratospherically derived air in a sample diminishes. In the INTEX A UT data set the range of values in HNO$_3$/O$_3$ that capture the clearly stratospheric samples also include many samples that could be in the tropospheric population, with HNO$_3$ < 500 ppt and O$_3$ < 120 ppb (Figure 6). In addition, processing in the troposphere (precipitation scavenging and photochemistry) will likely change the HNO$_3$/O$_3$ ratio over time.

The poor temporal and vertical coverage of $^7$Be data obtained in airborne campaigns is not likely to cause systematic over estimation of stratospheric influence. The technique used here assumes that all of the $^7$Be measured in the troposphere originated in the stratosphere, which is not strictly true. Subtracting a tropospheric fraction from measured $^7$Be would result in lower estimates of stratospheric influence, but it is not clear how to determine the partitioning between the two sources of $^7$Be. Likewise, we could require a minimum threshold of measured $^7$Be before using a data point to estimate the stratospheric contributions to simultaneously measured O$_3$ or HNO$_3$, but establishing that threshold would be arbitrary. Another possible flaw in the $^7$Be technique applied here (and previously, e.g., Dibb et al., 2000, 2003c) is in the treatment of the intercept derived from the regression analysis (Figure 3). It has been suggested that the intercept should be interpreted as the tropospheric background into which the stratospheric contribution is mixing (Daniel Jacob, personal communication, August, 2004). In this conceptual model, the stratospheric component of O$_3$ or HNO$_3$ would be given simply by measured $^7$Be x the slope found by regression, and the tropospheric residual would obviously be larger than the estimates presented in Figures 7 and 8.
We have applied this modified $^{7}\text{Be}$ approach to the INTEX A data set, to see if it would reconcile the different estimates of stratospheric influence on the composition of the UT based on HNO$_3$/O$_3$ and $^{7}\text{Be}$. Except for HNO$_3$ in the NE region above 8 km the two approaches are still divergent, with $^{7}\text{Be}$ suggesting a much stronger impact from STE (Figures 9 and 10). The estimated stratospheric contribution to UT O$_3$ above 8 km across the entire INTEX A region, based on this weaker link to $^{7}\text{Be}$, is 42% of the total, compared to the 10% contribution inferred from the HNO$_3$/O$_3$ ratio as a filter approach. For HNO$_3$ the comparison is 47% stratospheric contribution from $^{7}\text{Be}$ and 16% using the HNO$_3$/O$_3$ ratio.

**CONCLUSIONS**

The altitude distributions of $^{7}\text{Be}$, HNO$_3$, and O$_3$ above the INTEX A study region, and correlations between these species in the UT, provide evidence that stratospheric air was sampled occasionally during the mission. In addition, these three tracers suggest that air masses with stratospheric influence were encountered frequently above 6 km altitude, with distinct tropospheric and stratospherically influenced populations revealed by two, relatively compact, branches with different slopes in scatter plots of HNO$_3$ versus O$_3$ and HNO$_3$ versus $^{7}\text{Be}$. The slope of the stratospherically impacted branch of the HNO$_3$ – O$_3$ relationship in the UT during INTEX A was similar to that found in the lower stratosphere above northeastern North America during PAVE, suggesting that HNO$_3$/O$_3$ might serve as an additional tracer of stratospheric influence in the troposphere.

We have used the established $^{7}\text{Be}$ tracer, and HNO$_3$/O$_3$ as a proposed new tracer, to identify INTEX A samples in the UT with stratospheric influence, and to attempt to quantify the impact of STE on the mixing ratios of O$_3$ and HNO$_3$ in the UT. The two approaches both suggest that stratospherically impacted airmasses were frequently encountered, and that they were found most frequently in the northeastern portion of the INTEX A study region, in accord with other analyses of the INTEX A and ICARTT data sets.

Estimates of the fraction of HNO$_3$ and O$_3$ in the UT that was derived from the stratosphere were made by correlations to $^{7}\text{Be}$, and by using the HNO$_3$/O$_3$ ratio as a filter to exclude stratospherically impacted air masses. The approach using $^{7}\text{Be}$ suggests that
STE was the dominant source of both gases above 6 km altitude during the campaign, while the HNO$_3$/NO$_2$ filter indicates that STE is a significant (though not dominant) source of HNO$_3$ in the UT, but plays only a minor role in the UT O$_3$ budget. While the HNO$_3$/O$_3$ ratio does identify samples with strong (recent?) stratospheric influence, it appears to also select too many with little or no stratospheric influence, so that the subpopulation selected by this filter in the INTEX A data set has a tropospheric character. On the other hand, the O$_3$ and HNO$_3$ versus $^7$Be relationships overestimate stratospheric influence by not considering the presence of a small, but unknown, tropospheric fraction of $^7$Be in each sample. There is also uncertainty regarding whether the intercepts in the regression analyses should be included as part of the “stratospheric” HNO$_3$ and O$_3$, or considered the tropospheric endmember into which the stratospheric contribution is mixed. The latter approach yields smaller estimates of impact of STE on the UT, but even these estimates suggest that STE was a dominant factor in the budgets of UT HNO$_3$ and O$_3$ during INTEX A. It is likely that the true impact of STE on UT HNO$_3$ and O$_3$ lies between these two $^7$Be-based estimates. Measurements of additional stratospheric tracers, that were not available during INTEX A, were made during INTEX phase B in spring 2006; these may provide insight into the discrepancies between the different estimates of stratospheric influence on the composition of the UT described here.
REFERENCES CITED


Avery, M. A., et al. (this issue), Impact of multi-scale dynamical processes and mixing on the chemical composition of the upper troposphere and lower stratosphere as observed and modeled during INTEX A, submitted to J. Geophys. Res.


Bhandari, N., D. Lal, and Rama (1966), Stratospheric circulation studies based on natural and artificial radioactive tracer elements, Tellus, 18, 391-405.

Browell, E. V., M. A. Fenn, J. W. Hair, C. F. Butler, A. Notari, S. A. Kooi, S. Ismail, M. A. Avery, R. B. Pierce, and H. E. Fuelberg (this issue), Large-scale air mass characteristics observed during the summer over North America and Western Atlantic Ocean, submitted to J. Geophys. Res.


Table 1. Percentage of mist chamber sampling intervals with HNO$_3$/O$_3$ ratios characteristic of the lower stratosphere (2.0 – 4.0 ppt/ppb).

<table>
<thead>
<tr>
<th>Press. Alt. (km)</th>
<th>South West</th>
<th>North West</th>
<th>South East</th>
<th>North East</th>
</tr>
</thead>
<tbody>
<tr>
<td>6 – 12</td>
<td>34.8 (660)</td>
<td>30.4 (319)</td>
<td>30.5 (390)</td>
<td>42.4 (882)</td>
</tr>
<tr>
<td>10 – 12</td>
<td>37.4 (230)</td>
<td>10.8 (83)</td>
<td>17.4 (92)</td>
<td>45.4 (152)</td>
</tr>
<tr>
<td>8 – 10</td>
<td>34.6 (191)</td>
<td>28.7 (136)</td>
<td>27.5 (171)</td>
<td>42.5 (409)</td>
</tr>
<tr>
<td>6 – 8</td>
<td>32.6 (239)</td>
<td>49.0 (100)</td>
<td>44.1 (127)</td>
<td>40.8 (321)</td>
</tr>
</tbody>
</table>

1 (Total number of samples collected in each bin in brackets)
Figure 1. Example of the distribution of mistchamber and filter samples over time and altitude on a typical INTEX A flight.
Figure 2. Altitude profiles of $^7$Be, O$_3$, and HNO$_3$ from the DC-8 during INTEX A.
\[ O_3 = 50.9 + 0.039 \times {^7}\text{Be} \quad R^2 = 0.54 \]
\[ O_3 = 37.0 + 0.044 \times {^7}\text{Be} \quad R^2 = 0.58 \]

\[ \text{HNO}_3 = 85.7 + 0.118 \times {^7}\text{Be} \quad R^2 = 0.39 \]

Figure 3. Scatter plots of \( O_3 \) (upper panel) and \( \text{HNO}_3 \) (lower panel) versus \( ^7\text{Be} \). The solid least squares regression line in the \( O_3 \) plot uses all data points, while the dotted line applies \( ^7\text{Be} > 600 \text{ fCi m}^{-3} \) as a filter. For \( \text{HNO}_3 \) the fit is only for the strongly stratospherically impacted samples (\( ^7\text{Be} > 600 \text{ fCi m}^{-3} \)). Note that one boundary layer sample with \( \text{HNO}_3 > 8 \text{ ppb} \) is omitted from this plot for clarity.
Figure 4. Nitric acid versus ozone. In the upper panel all data points from INTEX A are shown in black diamonds on top of PAVE data (grey diamonds). The ratio of 3.5 ppt HNO$_3$/ppb O$_3$ was characteristic of measurement in the lower stratosphere during PAVE (Dibb et al., 2006). In the lower panel INTEX A observations above 6 km are plotted on an expanded scale to focus on the upper troposphere. The line has a 3.5 ppt/ppb slope, but intersects the x axis at 40 ppb O$_3$. 
Figure 5. Example time series (from INTEX A flight 11) of the relationships between HNO$_3$, O$_3$, and NO$_2$ in the upper troposphere. Only samples collected above 6 km are shown, but upper panel depicts the DC-8 flight profile. Grey bars highlight intervals with enhanced NO$_2$ from convection (Bertram et al, this issue), increasing the NO$_2$/HNO$_3$ ratio and depressing the HNO$_3$/O$_3$ ratio (lower panel). The mixing ratios of HNO$_3$ and O$_3$ outside of these convectively influenced intervals covary rapidly, but their ratio stays near 3.5 most of the time.
Figure 6. Nitric acid and ozone plotted against the HNO$_3$/O$_3$ ratio. Values of the ratio between 2.0 and 4.0 (grey shading) capture the stratospherically influenced HNO$_3$ and O$_3$ samples apparent in Figures 3 and 4. The stratospheric value of the HNO$_3$/O$_3$ ratio found during PAVE is shown by the vertical line.
Figure 7. Mean HNO$_3$ mixing ratios in four geographic regions and three upper tropospheric altitude bins sampled by the DC-8 in INTEX A. Black bar includes all data, dark grey bars exclude samples with recent convective influence (NO$_2$/HNO$_3$ >1.0) and the light grey bars exclude samples with stratospheric signature (2.0 < HNO$_3$/O$_3$ < 4.0).
Figure 8. As in Figure 7, but for $O_3$. 
Figure 9. Comparison of the HNO$_3$/O$_3$ filter estimates of stratospheric influence (dark grey bars, same as “minus stratospheric” in Figure 7), and the modified estimate using just the slope of the correlation to $^7$Be (light grey bars, slopes from Figure 3).
Figure 10. As in Figure 9, but for O$_3$. Dark grey bars are the same as in Figure 8.